



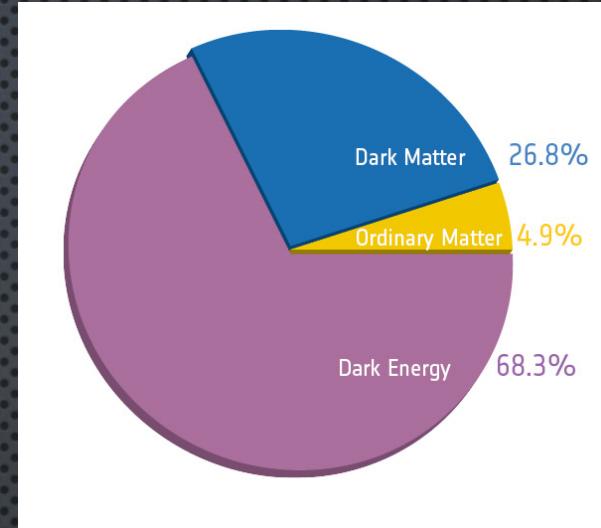
TOWARDS AN AB INTIO DESCRIPTION OF DARK MATTER SCATTERING

SAMUEL LEUTHEUSSER

WORK WITH S. RAGNAR STROBERG AND JASON D. HOLT

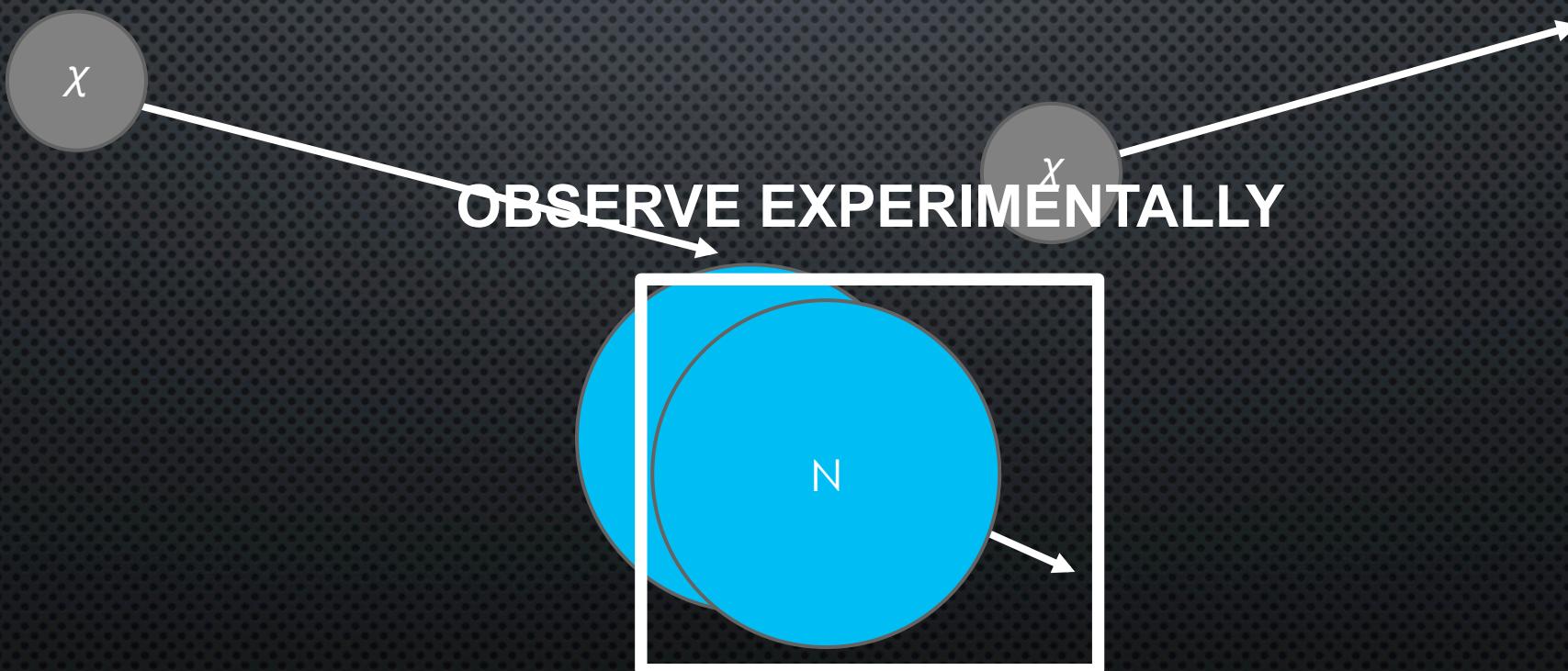
TUESDAY, FEBRUARY 28TH, 2017

DARK MATTER AS A PARTICLE



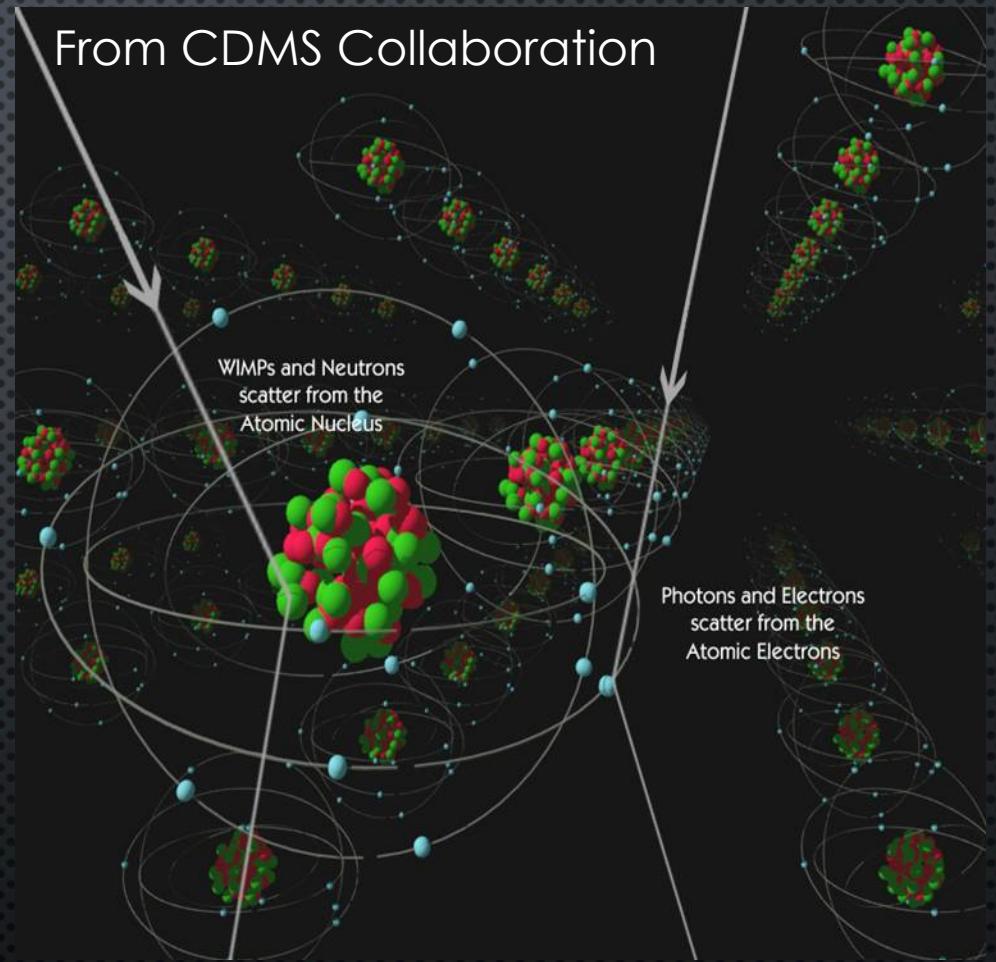
- COSMOLOGICAL OBSERVATIONS SUGGEST THAT OUR UNIVERSE IS COMPOSED OF ROUGHLY 27% DARK MATTER (DM)
- THESE OBSERVATIONS ARE OF THE GRAVITATIONAL EFFECTS OF DM
- DM IS NON-STANDARD MODEL MATTER WHICH INTERACTS VERY WEAKLY, IF AT ALL, WITH STANDARD MODEL PARTICLES
- MANY EXTENSIONS TO THE STANDARD MODEL PREDICT DM CANDIDATES CALLED WEAKLY INTERACTING MASSIVE PARTICLES (WIMPs)
- IF DM IS COMPOSED OF WIMPs, WE SHOULD BE ABLE TO SEE IT SCATTER OFF OF NUCLEI

DIRECT DETECTION EXPERIMENTS



DARK MATTER SEARCHES

- SUPERCDMS (NOW AT SNOLAB) – USES GE
- DEAP (SNOLAB) – USES AR
- DAMA/LIBRA (GRAN SASSO, ITALY) – USES NAI
- PICO (SNOLAB) – USES CCl_3F
- XENON (GRAN SASSO, ITALY) – USES XE
- LUX – USES XE
(SANFORD UNDERGROUND LAB, NORTH DAKOTA)



DESCRIBING DARK MATTER SCATTERING

- FOR WIMPS PREDICTED BY SUPERSYMMETRY, THE MOST IMPORTANT CONTRIBUTION TO SCATTERING COMES FROM AXIAL AND SCALAR CURRENTS
- THE AXIAL OR “SPIN-DEPENDENT” SCATTERING IS MUCH MORE SENSITIVE TO NUCLEAR STRUCTURE
- THE AXIAL STRUCTURE WAS GIVEN THE FOLLOWING DECOMPOSITION BY ENGEL, PITTEL AND VOGEL:

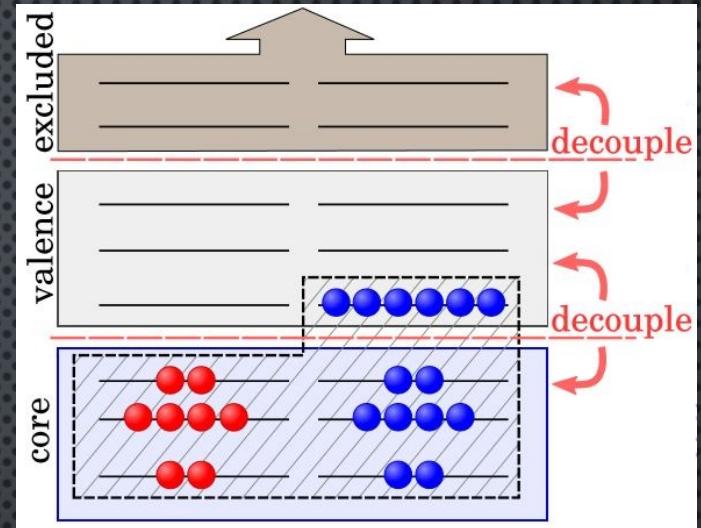
$$S \downarrow A(p) = \langle f | \mathcal{L} \uparrow A | i \rangle = a \downarrow 0 \gamma_2 S \downarrow 00(p) + a \downarrow 0 a \downarrow 1 S \downarrow 01(p) + a \downarrow 1 \gamma_2 S \downarrow 11(p)$$



Isoscalar Coupling

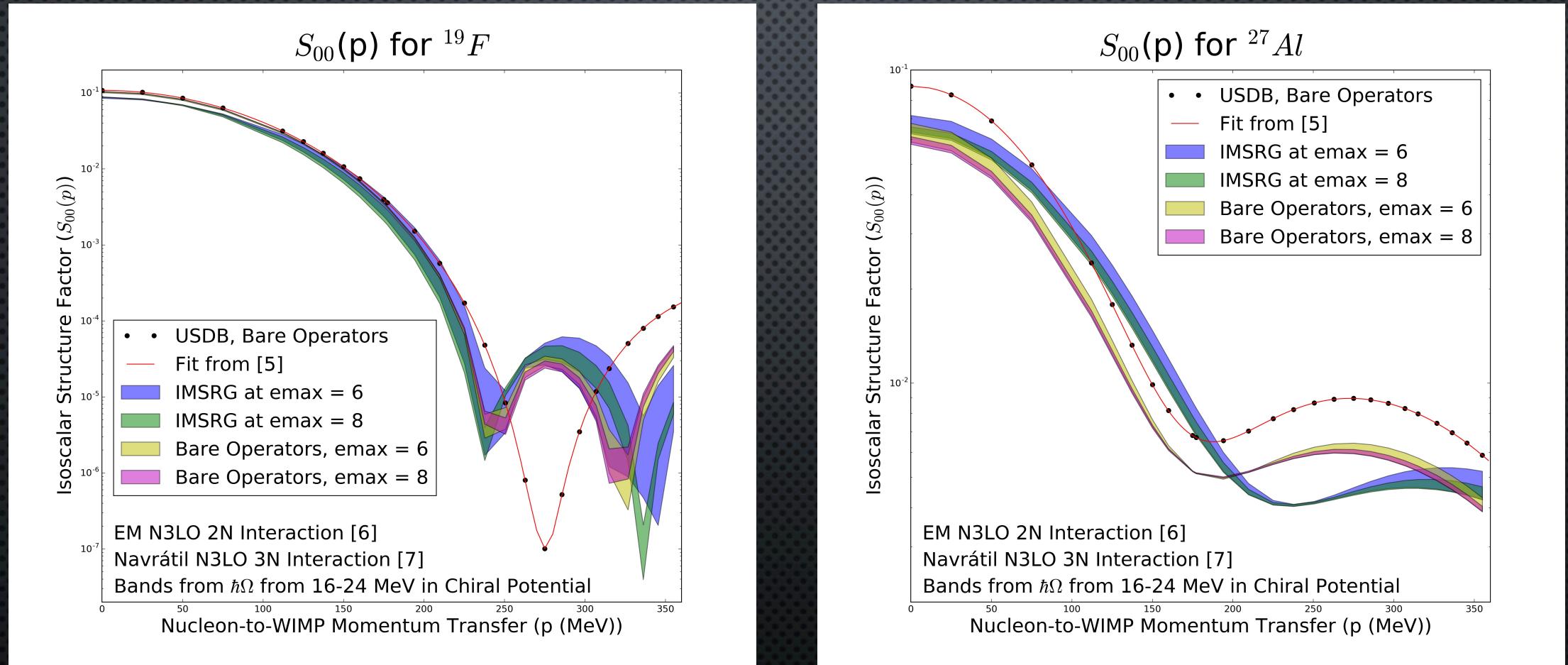
Isovector Coupling

VALENCE SPACE IMSRG



- FIND AN APPROXIMATE GROUND STATE USING HARTREE-FOCK AND A CHIRAL INTERACTION
- NORMAL ORDER ALL OPERATORS WITH THE HF STATE AS A REFERENCE
- PERFORM A UNITARY TRANSFORMATION DECOUPLING A VALENCE SPACE IN THE HAMILTONIAN
- SIMULTANEOUSLY TRANSFORM OTHER OPERATORS
- DIAGONALIZE THE VALENCE SPACE HAMILTONIAN FOR THE SPECTRUM
- COMPUTE MATRIX ELEMENTS USING TRANSFORMED OPERATORS AND NEW FOUND STATES
- THIS FRAMEWORK EASILY ALLOWS FOR CONSISTENT AB INITIO CALCULATIONS IN HEAVY NUCLEI

RESULTS



FUTURE DIRECTIONS

- COMPUTATION OF THE ISOVECTOR PIECES OF THE AXIAL STRUCTURE FACTOR
- INCLUSION OF FULL TWO-BODY TERMS
- STUDY OF UNCERTAINTIES FROM CHIRAL INTERACTIONS AND LOW ENERGY CONSTANTS
- CALCULATION IN HEAVIER NUCLEI: $^{129,131}\text{Xe}$, ^{127}I , ^{73}Ge